

2019/20  
**Master Mathematical Physics (Dijon) - General Relativity**  
**Exam**  
**8 juin (3 hours)**

*Every student commits, by participating in this exam, to respect the rules at the Université de Bourgogne, namely to answer without any external help to the posed questions.*

**COMPULSARY SIGNATURE :**

1. Given the metric on  $\mathbb{R}_+^2$ , defined by the line element ( $H$  is a constant):

$$ds^2 = \frac{H}{y^2} (-dt^2 + dy^2) ,$$

calculate the Ricci scalar  $R$ .

(Privilege the use of Cartan's structure equations; if not, use any method you know).

2. i) Consider (2-dim) Minkowski spacetime and timelike connected points  $p$  and  $q$ . Justify that a non-accelerated observer maximizes the proper time between  $p$  and  $q$ .
- ii) Justify that there exist observers travelling from  $p$  to  $q$  that make the elapsed proper time as small as desired (that is, there is no time-like curve that minimizes the proper time between  $p$  and  $q$ ).
- iii) Consider now Schwarzschild spacetime with mass  $M$  and two observers at fixed  $(\theta, \varphi)$ , initially at  $r_1 > 2M$ . At some given time, one of the observers descends to  $r_2$ , with  $2M < r_2 < r_1$ . After a very long time, this observer comes back to  $r_1$ . Who is older when the two observers meet? Compare with the situation in i): which physical principle can be invoked to establish an analogy between the situations i) and iii)?
- iv) We consider again these two observers in Schwarzschild. As before, one of them stays at a fixed  $(\theta, \varphi)$  at a given  $r$ , but now the second one follows a circular orbit without using the engines. Without making any calculation, just with a reasoning based on the understanding of i) and ii): which one will be older when they meet after one complete orbit?

3. Consider the spacetime  $\mathbb{R} \times ]-1, 1[$  with line element given by

$$ds^2 = -(1 - y^2)dt^2 + (1 - y^2)^{-1}dy^2 .$$

Change the line element from  $(t, y)$  to  $(\tau, y)$  coordinates using

$$\tau = t + \frac{1}{2} \ln(1 - y^2) .$$

In the  $(\tau, y)$  coordinates write explicitly the Klein-Gordon equation ( $m$  constant)

$$(\square - m^2) \phi = 0.$$

4. Given a time-orientable Lorentzian manifold, let us consider (at a given point) two vectors  $u^a$  and  $v^a$  which are timelike or null. If  $u^a$  is future directed show:
- i)  $g_{ab}u^av^b < 0$  if and only if  $v^a$  is future-oriented.
  - ii)  $g_{ab}u^av^b = 0$  if and only if  $u^a$  and  $v^a$  are null and collinear.
  - iii)  $g_{ab}u^av^b > 0$  if and only if  $v^a$  is past-oriented.